

Introduction

① Maxwell's equations

integral formdifferential form

Gauss

$$\oint d\vec{a} \cdot \vec{E} = \frac{Q}{\epsilon_0}$$

$$\rightarrow \vec{\nabla} \cdot \vec{E} = \frac{\rho}{\epsilon_0}$$

$$Q = \int d^3\vec{r} \rho(\vec{r})$$

$$\oint d\vec{a} \cdot \vec{B} = 0$$

$$\rightarrow \vec{\nabla} \cdot \vec{B} = 0$$

$$\phi_B = \int_S d\vec{a} \cdot \vec{B}$$

Faraday

$$\oint d\vec{l} \cdot \vec{E} = -\frac{\partial \phi_E}{\partial t}$$

$$\rightarrow \vec{\nabla} \times \vec{E} = -\frac{\partial \vec{B}}{\partial t}$$

$$\oint d\vec{l} \cdot \vec{B} = \mu_0 \epsilon_0 \frac{\partial \phi_E}{\partial t} + \mu_0 I$$

$$\rightarrow \vec{\nabla} \times \vec{B} = \mu_0 \epsilon_0 \frac{\partial \vec{E}}{\partial t} + \mu_0 \vec{J}$$

$$\phi_E = \int_S d\vec{a} \cdot \vec{E}$$

Ampere

$$I = \int_S d\vec{a} \cdot \vec{J}$$

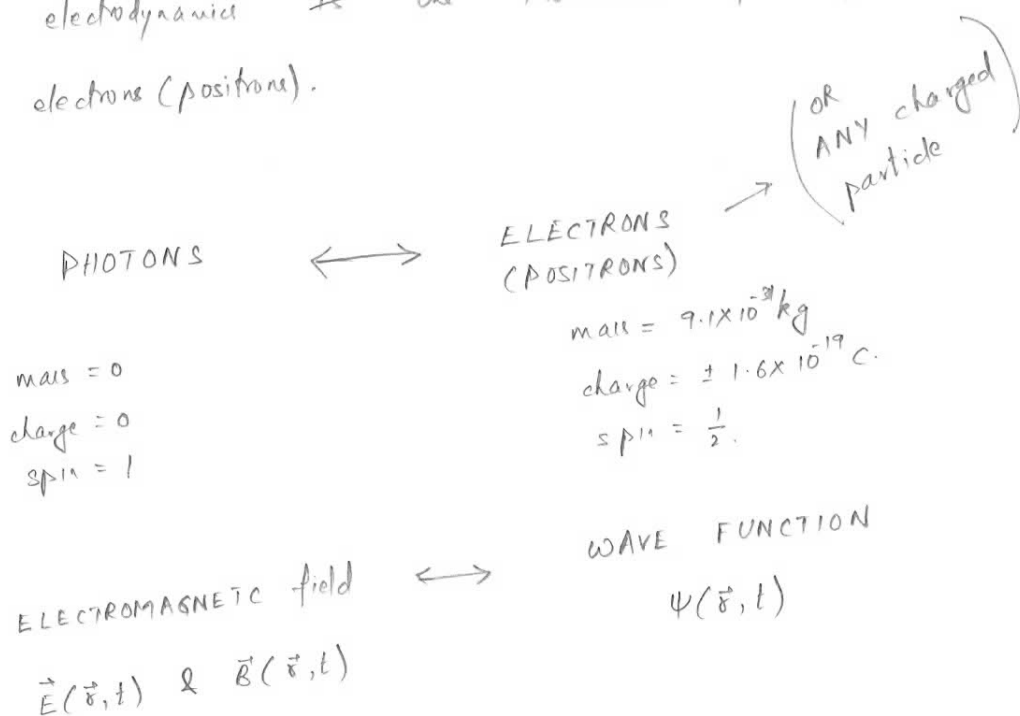
② Classical mechanics neglect $\left\{ \begin{array}{l} \text{(i) relativistic effects } \frac{v}{c} \ll 1 \\ \text{(ii) quantum effects } \Delta p \Delta x \gg \hbar \end{array} \right.$

→ space craft to moon — mostly classical mechanics.

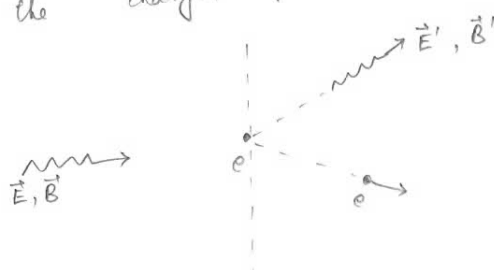
→ Light emission by atoms (LASER) — quantum mechanics.

→ pair production, Casimir effect } — quantum (relativistic) electrodynamics.

- ③ Quantum electrodynamics ~~is~~ describes the interaction of photons and electrons (positrons).



- ④ In this course we will neglect the dynamics of the charged particles.



Analogy: while studying the motion of a ball thrown by a boy on Earth, we can neglect the dynamics of Earth.

⑤ Examples of quantum electrodynamic phenomena that can not be understood by neglecting relativistic or quantum effects.

→ pair production by electric field

$$|\vec{E}_{\text{threshold}}| \sim 10^{16} \frac{\text{V}}{\text{cm}}$$

$|\vec{E}_{\text{threshold}}|$ for atmospheric breakdown is $10^4 \frac{\text{V}}{\text{cm}}$.

→ Casimir effect

two (neutral) metallic objects attract in vacuum.